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42 **Figure 1.** Tectonic setting around the 2015 Mw8.3 earthquake. Red dots: Epicenters of historical earthquakes near  
43 this Illapel 2015 event (Mw8.4 in 1922, Mw8.3 in 1943, Mw7.8 in 1971, Mw7.8 in 1985, Mw8.8 in 2010). Yellow  
44 dot: Epicenter of 2015 Mw8.3 Illapel event. Black circles: aftershocks (from <http://earthquake.usgs.gov>, as of  
45 18/9/2015). The red barbed line is the Chile trench trace. Color stripes along the trench depict past earthquake rupture  
46 zones (adapted from Vigny et al., 2011). ETOPO1 Digital Elevation Models  
47 (<http://www.ngdc.noaa.gov/mgg/global/global.html>) were used to generate the background topography. The black  
48 rectangle is the fault plane projected onto the surface.

49  
50 The information on the down-dip limit of the seismogenic zone and transition depth from seismic to aseismic  
51 slip of thrust faulting earthquakes is important to understand Chile subduction zone. (Mendoza et al. 1994; Pritchard  
52 et al. 2006). Modern geodetic technology can obtain small deformation of crust and could be used as a tool for  
53 seismic hazard assessment (Ader et al. 2012). Inversion of the co-earthquake rupture depth constrained by a dense  
54 geodetic data, e.g., deformation measurement from Synthetic Aperture Radar Interferometry (InSAR) and Global  
55 Positioning System, permits to address this issue. For example, using InSAR and GPS data, Tong et al. (2010)  
56 estimated the maximum rupturing depth of the 2010 Chile Mw8.8 event, which is 43-48km and is largely consistent  
57 along the 600km-long rupture zone. For the same event and also using joint inversion of ALSO/PALSAR and GPS  
58 data, Pollitz et al. (2011) suggested that the fault rupture of this event terminated at a depth of 35km, which is relative  
59 shallow, and likely associated with the spherical layering Earth model used in their inversion. Using joint inversion  
60 of teleseismic records, InSAR and high rate GPS (HRGPS) data, Delouis et al. (2010) constrained the maximum  
61 down-dip depth as 50km for the 2010 Chile great shock. These studies on the rupturing depth of the great earthquakes  
62 can provide evidence for determining the seismogenic depth, lower limit of stick-slip and the boundary between  
63 seismic and aseismic layers in the subduction zone beneath central Chile.

64 In this work, both the descending and ascending track of Sentinel-1A/IW data have been downloaded and  
65 processed to reconstruct the coseismic deformation field of the 2015 Chile event. Then, we inverted the slip  
66 distribution on the seismogenic fault plane of this earthquake with three different constraints of descending and  
67 ascending measurements separately and jointly. Thirdly, we discussed the observation and inversion results.



68 Additionally, inversion results (e.g., rupture depth) from each track (ascending, descending) and join tracks has been  
 69 analyzed. Finally, in order to identify the promoting relationship between the main shock and aftershocks, we  
 70 estimated the shear stress in the aftershock area of the main shock event. The complete InSAR coverage over the  
 71 rupture area provided a unique information to derive a detailed slip model, which is needed to estimate the spatially  
 72 varied stress change from the event.

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## 74 2. Sentinel-1A InSAR data and processing

75 We investigated the crustal deformation triggered by the 2015 Mw8.3 Chile earthquake using interferometric  
 76 synthetic aperture radar (InSAR) with Sentinel-1A Interferometric Wide Swath (IW) mode data in both descending  
 77 and ascending orbits. Sentinel-1A satellite was launched by ESA on April 3, 2014, and its IW Mode used the  
 78 advanced TOPSAR (Terrain Observation with Progressive Scans SAR) technology  
 79 (<https://sentinel.esa.int/web/sentinel/user-guides/sentinel-1-sar/applications>). The radar image in IW model has a  
 80 swath width up to 250km, spatial resolution of 5m×20m (single look), and revisit period of 12 days, providing a  
 81 good data source for large-scale monitoring of ground deformation. At present, Sentinel-1A satellite data can be  
 82 accessed through ESA data hub (<https://scihub.esa.int/>). It is one and three days after the 16 September 2015 Chile  
 83 event, i.e. 17 September and 19 September that Sentinel-1A acquired descending and ascending data covering the  
 84 coseismic area. The selected post-earthquake images are close to the event time, whereas the pre-image can be  
 85 acquired long before the event. Because data very close to the mainshock time, permit to study the coseismic  
 86 deformation of this event without much aftershock deformation. As the affected area of this great event is very large,  
 87 we use three adjacent frame along the same descending track to get a full coseismic deformation field. Since, in the  
 88 ascending track, only two frames are available, we get only part of coseismic deformation field. The SAR data and  
 89 its parameters used in this paper are shown in Table 1.

90

**Table 1.** The Sentinel-1A/IW data used in this study

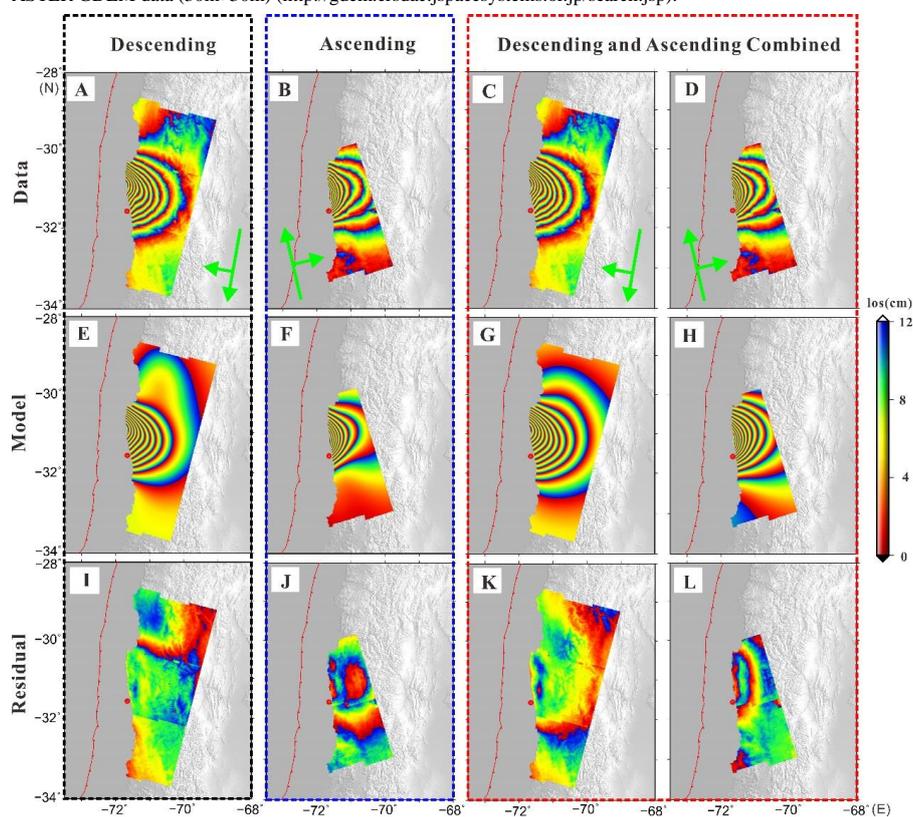
Number	Track	Master	Slave	Average Perpendicular Baseline(m)	Average Ambiguity Height(m)
1	Descending	20150707	20150917(north)	1	13667
2	Descending	20150707	20150917(middle)	-1	13667
3	Descending	20150707	20150917(south)	-3	4556
4	Ascending	20150826	20150919(north)	73	187
5	Ascending	20150826	20150919(south)	70	195

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92 We used the GAMMA software to process the Sentinel-1A data. The interferograms have been processed  
 93 separately for each frame along the same track, and then mosaicked to a signal wrapped differential interferogram.  
 94 To reduce noise, multi-look processing of 10-sight in range and 2-sight in azimuth directions were performed to the  
 95 interferograms. It requires a very high SLC (single look complex image) registration accuracy in azimuth direction  
 96 (Meta et al., 2010). To achieve an accuracy of a very small fraction of an SLC pixel nearly 0.1% in azimuth direction,  
 97 we performed intensity image based and iterating offset estimation for many times until the azimuth offset correction  
 98 became at least smaller than 0.02 SLC pixel. Meanwhile, the adaptive filters based on interferometric fringe  
 99 frequency and gradually decreasing windows were applied to interferograms so that their ratio of signal to noise was  
 100 highly enhanced, fringes associated with seismic deformation were highlighted. The algorithm of minimum cost  
 101 flow (Werner et al., 2002) was implemented for phase unwrapping with Delaunay triangle network that is suitable



102 for low coherent areas. To make phase continuous and smooth, before integration of mosaic three adjacent  
 103 interferograms on descending track, we firstly unwrapped the interferogram in the southernmost of the study area,  
 104 and used the far field to its south as the start point for unwrapping. It was followed by unwrapping the interferogram  
 105 in the middle, using the same-place point in superposed portion of the two adjacent interferograms as the reference  
 106 and initial phase value for unwrapping it. Similarly, the interferogram in the north was unwrapped. Consequently,  
 107 the interferogram from integrating these three images was featured by continuous phase without signature of  
 108 boundaries. When doing this, we removed topographic phase by the generating simulated interferogram using  
 109 ASTER GDEM data (30m×30m) (<http://gdem.ersdac.jspacesystems.or.jp/search.jsp>).



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 111 **Figure 2.** Earthquake mainshock, as seen from radar satellites that allow quantifying displacements in the Line of  
 112 Sight (LOS) as indicated by green arrow in (A-D). Upper row is the data (A, C, coseismic deformation field from  
 113 descending track; B, D, coseismic deformation field from ascending track). Center row (E-H) is the model solution  
 114 (the model in black dashed area is constrained by descending data, in blue dashed area is constrained by ascending  
 115 data, in red dashed line is constrained by ascending and descending data combine). Bottom row (I-L) shows the  
 116 residual after subtracting the model from the data. Red dot depicts the location of the mainshock epicenter from  
 117 USGS, about 10km to the coast. The red barbed line is the trench trace.

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 119 **3. Coseismic deformation fields derived from Sentinel-1A ascending and descending InSAR data**

120 As mentioned above, the deformation field from descending data was generated by integrating three  
 121 interferograms along the same track, which covers almost the whole affected area of the 2015 Chile earthquake (Fig.  
 122 2A). Towards the continent, the fringes become progressively sparse, implying decreasing gradients of deformation.



123 While we set the deformation in far field, without any phase change, to zero, the maximum LOS displacement is -  
124 133cm near the coast. It looks like a half circle convex to east, with most of data being negative, which means  
125 subsidence in descending LOS direction. According to the full descending track fringes, the deformation area is  
126 within 300km long in the NS direction or along the coast, and 190km in the EW direction. Although the deformation  
127 field derived from the ascending data is only based on two frame, it also covers the major part of the seismic  
128 deformation area (Fig 2B), consistent with that from the descending data. This field is also of a half circle convex to  
129 east with maximum LOS displacement 132cm (far field deformation be zero). Deformation in LOS being positive,  
130 means uplift in ascending LOS direction. The positive and negative with similar magnitude of LOS deformation  
131 from ascending and descending data suggest that the crustal deformation caused by this earthquake is dominated by  
132 horizontal motion.

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#### 134 4. Fault slip inversion and interferogram simulation

##### 135 4.1 Inversion method and fault model construction

136 Focal mechanism solutions given by USGS (<http://www.usgs.gov/>) and GCMT  
137 (<http://www.globalcmt.org/CMTsearch.html>) show that the seismogenic fault of this earthquake is thrust with a small  
138 dip angle. Its surface trace closely follows the trench axis. Based on the focal mechanism solutions, aftershock  
139 distribution and InSAR deformation fields obtained in this work, we built a single-plane fault model in elastic half-  
140 space (Okada Y. 1985) to invert the static coseismic slip distribution on the rupture surface constrained by the  
141 Sentinel-1A descending and ascending data both separately and jointly. The linear-inversion, Sensitivity Based  
142 Iterative Fitting (SBIF) method (Wang et al., 2008) was employed. Firstly, the fault plane was divided into multiple  
143 fault patches. Each patch was presumed to slip uniformly. In this way, the non-linear problem can be transformed  
144 into a linear problem. Then we used the mean square deviation reducing function to quantify the misfit between the  
145 simulated interferogram and the observed one. Using this function, by minimizing mean square deviation, non-  
146 uniform slip distribution on the fault plane can be determined. The mathematical formulation of the inversion is  
147 expressed by:

$$148 f(s) = \sum_{k=0}^K \|D_k - D_k^0 - G_k s(x)\|^2 + \beta^2 \|Hs\|^2 \rightarrow \min$$

149 Where 's(x)' is slip vector, 'k' is the patch index of different input data set. 'D' is the matrix of observation data, 'D<sup>0</sup>'  
150 is the static offset of the observations, 'G' is the Green's function for an elastic half-space, which describes the  
151 relation between the model prediction and the observation. 'β' is defined as the smoothing factor, 'H' is the Laplacian  
152 operator, and  $\|Hs\|^2$  represents the slip roughness. Assuming a Poisson ratio of 0.25 and using SBIF program, we  
153 calculate the Green's functions of the homogeneous elastic half-space model using Okada.

154 Also, we resampled the InSAR deformation field by the quad-tree resampling method for inversion (Jónsson  
155 et al., 1999; Lohman et al., 2005). The reason to do quad-tree is to reduce computation load and also to keep the  
156 pattern of deformation map. In the resampling process, we have 12763 sampled points from Sentinel-1A descending  
157 data and 9196 sampled points from Sentinel-1A ascending data, respectively, which still have a much higher spatial  
158 density than other geodetic data (e.g., GPS). The initial fault geometry is a single planar surface striking N4.6°E and  
159 dipping toward the east, where it takes trial values between 10° and 30°. The rake angles are in the range of 80°-150°.  
160 The upper boundary of the fault is to surface. The initial fault model is steadily modified through optimal fitting to  
161 the deformation fields derived from both Sentinel-1A descending and ascending data sets in joint inversion. The  
162 final fault model is that dip angle is 18.3°, strike is 4.6°, and the fault dimensions are 535 km along-strike and  
163 200 km down-dip, with fault plane divided into many rectangular patches whose grid is 10km×10km from best fault  
164 resolution test (Table 2).



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**Table 2.** Main parameters of the optimal fault model

Parameters	Lat_Ref(°)	Lon_Ref(°)	Strike(°)	Dip(°)	Length(km)	Rake(°)	Width(km)	Top_Depth(km)
Final values	-33.3	-72.75	4.6	18.3	535	80~150	200	0

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#### 4.2 Fault slip inverted from Sentinel-1A descending and ascending data separately and jointly

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Using the fault model and resampled data points described above, we inverted slip distribution on the fault plane of this Chile earthquake constrained by the Sentinel-1A descending and ascending observations separately and jointly. The calculated residual maps between the observed and simulated ones are shown in Figure 2. The result shows that when the inversion is constrained by Sentinel-1A descending data alone, the preferred slip model shows a preponderant fault rupture zone located in the shallow part of the up-dipping thrust fault above the hypocenter (Fig. 3B). The maximum fault slip is over 8 m at a shallow depth, located in the northwest of the epicenter. The down-dip boundary of the rupture zone is relatively clear, and its depth is only about 35km under the surface. The rupture length of the slip area is about to 340km, comparable to 335 km of the major axis of aftershock distribution in north and south direction. But the main slip is concentrated in a shallow region that is 15km deep and 200 km long on the subduction interface. The mean rake angle from inversion is 110°, consistent with the thrust fault motion. The simulated interferogram is reconciled well with the observed one, with fitting degree 99.99%. The seismic moment magnitude is Mw8.27.

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When the inversion is constrained by the Sentinel-1A ascending data alone, the resulted fault slip magnitude and its scope are all smaller than that constrained by the Sentinel-1A descending data, although the overall patterns of the slip region in both cases are similar (Fig. 3A). The probable reason is that the ascending data do not fully cover the deformation field, since only two frame images are available. The maximum slip from this inversion is only about 3.43m. The mean rake angle is 102.42°. The simulated interferogram fits the observed one very well with the fitting degree 99.97%. The seismic moment magnitude is Mw8.09. It should be pointed out that the down-dip boundary of the rupture zone is much deeper, with the depth about 50km under the surface.

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**Table 3.** Fault plane and source parameters of the 2015 Chile earthquake given by teleseismic focal mechanism solutions and this study

Source	Latitude (°)	Longitude (°)	Depth (km)	Mw	Strike (°)	Dip (°)	Rake (°)	Scalar Moment(N.m)
GCMT	-31.22	-72.27	17.8	8.2	5	22	106	$2.455 \times 10^{21}$
USGS	-31.57	-71.67	20.7	8.3	5	22	106	$3.467 \times 10^{21}$
Descending	-	-	-	8.27	4.6	18.3	116.41	$3.126 \times 10^{21}$
Ascending	-	-	-	8.09	4.6	18.3	102.42	$1.679 \times 10^{21}$
Jointly	-	-	-	8.25	4.6	18.3	103.24	$2.917 \times 10^{21}$

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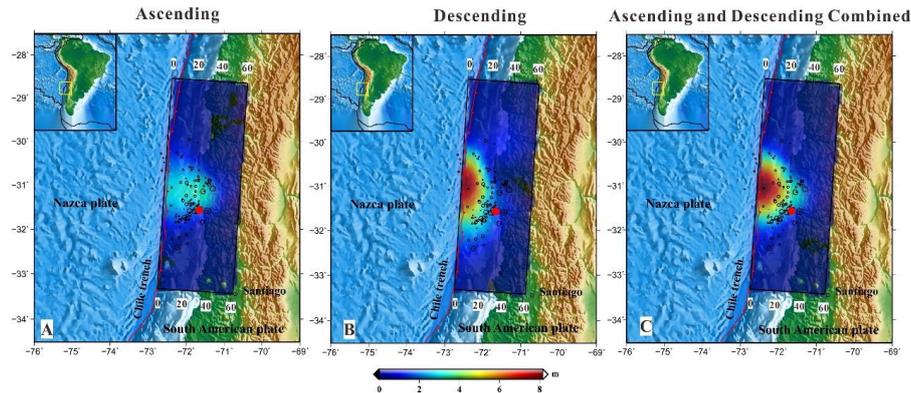
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We implemented a fault slip inversion jointly using the Sentinel-1A ascending and descending data with equal weight. The result falls between the two inversion results by using the two data alone (Fig. 3C). The shape of slip area seems to be symmetrical. The inversion result indicates that the mean rake angle is about 103.24°, which is in agreement with a thrust fault. The fitting degree is also very good, about 99.97%. The maximum slip is about 8.16m at shallow depth near the trench. The seismic moment magnitude is Mw8.25. The final inversion results are shown in Table 3 and Figure 3. It seems that combination of descending and ascending InSAR data used in inversion helps to derive a more comprehensive fault slip distribution.



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200 **Figure 3.** Fault slip distribution inverted by using Sentinel-1A ascending and descending data. The fault trace at the  
 201 surface is from (-33.3N, -72.75E) in south to (-28.5N, -72.30E) in north, strike is N4.6 E. The blue rectangle is the  
 202 fault plane projected onto the surface. Number with white background is the depth of the fault in kilometer. Red dot  
 203 is the position of epicenter from USGS. Aftershocks (from <http://earthquake.usgs.gov>, as of 18/9/2015) are  
 204 represented by black circle. (A) Fault slip distribution inverted by ascending data. (B) Fault slip distribution inverted  
 205 by descending data. (C) Fault slip distribution inverted by ascending and descending data jointly as constraints.

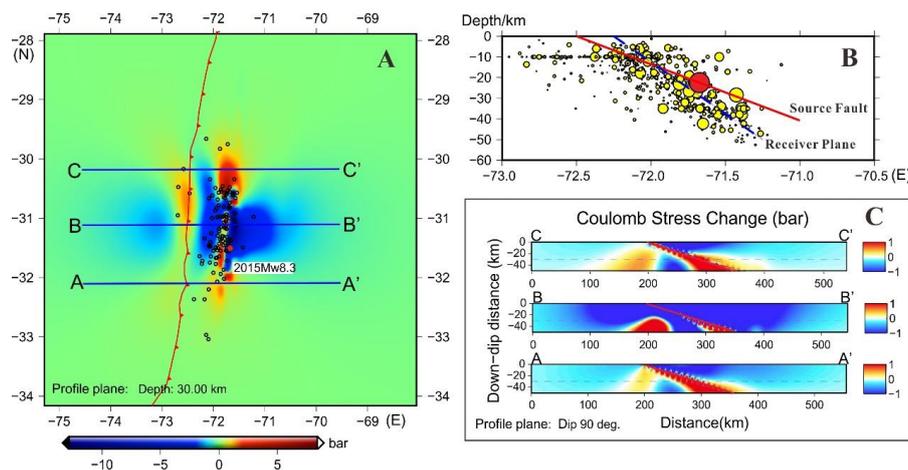
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#### 207 4.3 Static Coulomb stress changes

208 In order to identify the promoting relationship between the main shock and aftershocks, we calculated the  
 209 coseismic Coulomb Failure Stress (CFS) change on the fault plane and surrounding medium by using the optimal  
 210 slip model (Lin 2004; Toda 2005), which comes from inversion by ascending and descending jointly. Computing the  
 211 CFS change following an earthquake tells whether a fault has been brought closer or away from rupture (Stain 1999).  
 212 Many researches suggest that 0.1bar shear stress change can have a great influence in earthquake activities (King et  
 213 al., 1994). From the location of aftershocks, we found most of the aftershocks happened under our inverted fault  
 214 plane (Fig. 4B red line). The distribution of aftershocks reflected a special plane (Fig. 4B dash line in blue) whose  
 215 dip angle is about  $31^\circ$ , much bigger than the dip angle ( $18.3^\circ$ ) from our main shock inversion and results from USGS  
 216 (dip= $19^\circ$ ) and GCMT (dip= $22^\circ$ ). So in our models we set the receiver plane dip angle  $31^\circ$ , with strike  $4.6^\circ$  and rake  
 217  $105^\circ$ , to see what the static coulomb stress the main shock promoted to aftershock is. From the coseismic Coulomb  
 218 stress profile at 30km depth (Fig. 4A), we estimate the coseismic shear stress change ranged from -12 bar (stress  
 219 drop) to 8 bar (stress increase) and find aftershocks (depth in 20km-30km) locations correlate well with the areas  
 220 having increased Coulomb stress. The three special profiles in vertical (dip= $90^\circ$ ) reflected most areas with increased  
 221 Coulomb stress appeared beneath the main shock fault plane, which is consistent with the location where aftershocks  
 222 took place. At the same time, we can see static Coulomb stress up the main shock fault plane is released (Fig. 4C).  
 223 A frictional ratio of 0.4 and rake angle 105 were used in these results, but we also explored different frictional ratios  
 224 (0.3-0.7) and rake angles of receiver plane ( $100^\circ$ - $110^\circ$ ). No significant difference was observed for the obtained CFS,  
 225 implying that the models are robust.

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 229 **Figure 4.** (A) Coulomb Failure Stress (CFS) changes (in bar) due to 2015 September 16 main shock is calculated at  
 230 a depth of 30km. The red circle marks the main shock and black circle is the location of other aftershocks (Mw>4)  
 231 in the depth from 20km to 30km. (B) Location of aftershocks (from <http://earthquake.usgs.gov>, as of 11/10/2015) in  
 232 vertical plane. The main shock is the red dot and the red line is the fault plane in our inversion whose dip angle is  
 233 18.3°. Blue dash line is the receiver plane which is best-fit from the location of aftershocks. The dip angle of the  
 234 receiver plane is about 31°. (C) Cross section with the main fault lines (red line) and CFS calculated by the coseismic  
 235 mainshock slip model, where red denotes a stress built up, and blue a stress shadow. The areas beneath the fault  
 236 plane receive large CFS built up.

## 238 5. Discussion

239 This paper presents a study of 2015 Chile Mw8.3 earthquake based on Sentinel-1A InSAR data of ascending  
 240 and descending tracks. The purpose is to investigate the coseismic deformation and invert the slip distribution on its  
 241 fault plane and rupture depth. The results show that the overall slip area is located in the shallow portion of the  
 242 subduction interface between the source and the trench, with a NS symmetric pattern. The moment magnitude (8.25)  
 243 and seismic moment ( $2.917 \times 10^{21}$  Nm) are between the results of GCMT ( $M_w=8.2$ ,  $M_0=2.455 \times 10^{21}$ ) and USGS  
 244 ( $M_w=8.3$ ,  $M_0=3.467 \times 10^{21}$ ), consistent with focal mechanism solutions from seismic waves. It indicates that the  
 245 inversion results of this work are reliable, including the dip angle 18.3° which is consistent with result from USGS  
 246 (dip=19°) and GCMT (dip=22°). Fault slip is likely related with the ground broader deformation field in EW  
 247 direction from our work. To assess the resolution capabilities and stability of the fault model, we conducted fault  
 248 resolution tests of slip identification and find sub-fault grid 10km×10km is best.

249 The coseismic deformation fields of the Chile event derived from Sentinel-1A descending and ascending data  
 250 are roughly consistent in the shape. The positive and negative with similar magnitude of LOS deformations from  
 251 ascending and descending data suggest that the crustal deformation caused by this earthquake is dominated by  
 252 horizontal motion. The fault-slip distributions of the Chile event from inversions constrained by different data sets  
 253 (Sentinel-1A descending, Sentinel-1A ascending, and jointly) have both similarities and disparities. The common  
 254 points include that the mean rake angles (102°-106°), indicating a thrust fault with slight right lateral slip, and the  
 255 outlines of slip regions are about in the same area from the three inversions. There are differences between the  
 256 estimated slip (~3m) from inversion using Sentinel-1A ascending data alone and slip (~8m) from inversion using  
 257 Sentinel-1A descending data alone or inversion jointly. An alternative explanation is that the ascending displacement  
 258 field is smaller than the actual one because two-frame images cannot cover the whole coseismic deformation field.



259 With reference to all the results, we suggest that the slip-distribution from the inversion using Sentinel-1A descending  
260 and ascending data jointly seems to be more convincing.

261 Here we also compare the surface deformation fields and slip distributions on the fault planes of the 2015  
262 Mw8.3 and 2010 Mw8.8 Chile events. Although the South American subduction zone hosts a significant number of  
263 large earthquakes, only these two events have InSAR data available for such a comparison. InSAR data, owing to  
264 the advantages of dense sampling, can provide the best constraint on the slip location, distribution and depth on the  
265 rupture plane by quantitative measuring the static displacement on the ground surface caused by an earthquake. We  
266 find that the two events are different in coseismic deformation. For the 2010 event, the deformation spreads along  
267 the coast with at least two centers (Tong et al., 2010; Bertrand et al, 2010). The slip distribution from inversion is  
268 also of a narrow long strip, rupturing over 600km. The slip is concentrated on the north and south of the source,  
269 mostly at depths of 15km-25km, and no large slip at the trench. In contrast, the deformation field of the 2015 event  
270 is a complete half circle shape. The inverted slip concentration area is nearly NS symmetric, close to the shallow  
271 trench. The 2015 event is located over 400km north of the 2010 shock, both on the subduction slab of the Nazca  
272 plate beneath the South American plate. Both events are interplate thrusts with similar tectonic and dynamic settings.  
273 But as mentioned above their rupture features are different. The analysis suggests that the 2015 event has a shallow  
274 source (25km) and a connective rupture in up-dip direction above the source reaching the trench. Meanwhile its main  
275 shock occurred on a big barrier. While the 2010 event is relatively deeper (33km), and at least ruptured two big  
276 barriers. It may induce to speculate that the subduction zone has many barriers of varied sizes on different segments.  
277 And the coupling or locking degrees are variable at different sections of the subduction zone.

278 However, the maximum rupture depth (50km) of the 2015 Mw8.3 event from the model of this work is roughly  
279 consistent with the rupture depth of the 2010 Mw8.8 shock derived from inversion of previous studies, which are  
280 based on InSAR plus GPS or InSAR, GPS, and seismic wave data (Tong et al., 2010; Bertrand et al, 2010). It is also  
281 in accordance with the depths of the subduction zone in northern Chile (Tichelaar et al., 1993; Delouis et al. 1997)  
282 and southern Chile (Delouis et al. 2009) and the locking depth of this zone from GPS data (Ruegg et al., 2009).  
283 Previous studies suggest that this depth is the transition between the seismic and aseismic layer in the subduction  
284 zone beneath the South American plate. This work further confirms this conclusion.

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## 286 6 Conclusions

287 In this work, we obtained the coseismic deformation field of the 2015 Chile Mw8.3 earthquake using Sentinel-  
288 1A descending and ascending data. The positive and negative with similar magnitude of LOS deformation from  
289 ascending and descending data suggest that the crustal deformation is dominated by horizontal motion. The inversion  
290 constrained by Sentinel-1A ascending and descending data jointly can display comprehensive fault slip. We find the  
291 strike angle  $N4.6^\circ E$ , the fault dimension 535km (along-strike)  $\times$  200km (down-dip), and the dip angle  $18.3^\circ$  can fit  
292 our model better. Mean rake angle from inversion is  $103.24^\circ$ , which indicates a thrust fault with slight right-lateral  
293 slip. A maximum slip of 8.16m on the fault plane appears near the trench, the length of rupture reaches about 340km  
294 along strike but mainly extending to north side of the epicenter, and the overall slip pattern is moderately symmetrical,  
295 with the down-dip end of the rupture at about 50km which is roughly consistent with the rupture depth of the 2010  
296 Mw8.8 shock. The seismic moment magnitude is Mw8.25, the scalar moment from jointly inversion is  
297  $(2.917 \times 10^{21} \text{ N m})$ , and the fitting degree of the whole field is 99.97%. Coseismic Coulomb stress change reflected  
298 most areas with increased Coulomb stress appeared beneath the main shock fault plane, which is consistent with the  
299 location where aftershocks took place. At the same time, we can see static Coulomb stress above the main shock  
300 fault plane is released.

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306 All figures were generated with Generic Mapping Tools software.

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